



An integrated view on high density operation and fuel cycle

Part I

R. Wolf, C. Angioni, M. Bernert, C. Day, A. Kallenbach, V. Kotov, P. Lang, M. Maraschek, D. Reiter, G. Tardini, H. Zohm, et al.



robert.wolf@ipp.mpg.de



Contents



- Motivation
- Density limits
- Tokamak / stellarator
 - Radiation limit
 - Greenwald density as pedestal limit
 - Pellet fuelling
 - Divertor neutral particle pressure
 - Divertor heat flux
- Summary / workplan

Motivation



High density operation is desirable for maximizing fusion power

$$P_f = \int n^2 < \sigma v > E_f dV$$

at $T_i = 10 - 15 \text{ keV}$

$$\sigma v >= c_{\sigma v} T^2$$

with

$$B = \frac{p}{B^2/2\mu_0}$$

<

β

$$P_f = 2\mu_0 c_{\sigma v} E_f \int \beta^2 B^4 dV$$

- thus, below β -limit increase of density in order to stay in the optimum fusion reaction range (at 10 15 keV)
- Cost of electricity (in tokamaks) drops with increasing n/n_{GW}

$$coe \propto \left(\frac{1}{A}\right)^{0.6} \frac{1}{\eta^{0.5}} \frac{1}{P_e^{0.5} \beta_N^{0.4} \left(\frac{n}{n_{GW}}\right)^{0.3}}$$

D. Ward et al.

Motivation



In addition

• decrease of population of fast particles (of particular importance for stellarators)

$$T_s \propto \frac{T_e^{3/2}}{n_e} \propto \frac{p_e^{3/2}}{n_e^{5/2}}$$

- avoid fast particle driven instabilities altogether?

 $\mathbf{2}$

- for stellarators high density required for high Q_{DT} or ignition, because of

$$D \sim \varepsilon_{eff} T^{7/2} = \varepsilon_{eff} \left(\frac{p}{n}\right)^{7/2}$$

- improve lifetime of plasma facing components / divertor; higher densities / lower temperatures
 - reduce local heat fluxes due to increased radiation / detachment
 - reduce high energy ion tail impinging on target
- relation between plasma density, neutral density, plasma fuelling and fuel cycle
 - maximize core density, divertor density and neutral density at the pumps
 - minimize tritium turn-over

Density limits: tokamaks and stellarators

• In tokamaks density empirical density limit given by Greenwald scaling

$$n_{GW}[10^{20}m^{-3}] = \frac{I_p[MA]}{\pi a^2[m^2]}$$

- limit associated with loss of confinement (H to L transition)
- disruptions
- In stellarators density limit
 - no Greenwald limit (e.g. LHD super dense core plasma, up to $5 \times n_{GW}$)
 - theoretically given by bremsstrahlung
 - in practise often radiation limit is observed (Sudo-limit)

 $\bar{n}_{Sudo} \propto rac{P \cdot B}{V}^{0.5}$ associated with impurity accumulation

Workshop on MFE Roadmapping, Princeton 7-10 September

6

Comparison stellarator and tokamak

Tokamaks:

- Temperature screening counteracts effect of density peaking
- Density peaking increases with decreasing collisionality, also considering central heating by fusion α's (next slide)

Stellarators

- Thermodynamic forces are predicted to support accumulation in the standard case with negative radial electric field, the socalled ion-root regime
- Suitable confinement regimes have to be found (→ HDH)
- Test of high-Z (tungsten) still has to be done



7

Comparison stellarator and tokamak

Tokamaks:

- Temperature screening counteracts effect of density peaking
- Density peaking increases with decreasing collisionality, also considering central heating by fusion α's (next slide)

Stellarators

- Thermodynamic forces are predicted to support accumulation in the standard case with negative radial electric field, the socalled ion-root regime
- Suitable confinement regimes have to be found (→ HDH)
- Test of high-Z (tungsten) still has to be done

HDH mode in W7-AS characterized by steep edge ∇n





Tokamak: density peaking scales with collisionality



Effect of density peaking for a tokamak DEMO

- Hypothesis: Greenwald limit is a pedestal limit
- ASTRA simulation of DEMO (D. Ward: R/a = 8.5/2.83, 5.74 T, 23 MA, $\beta_{N,th}$ = 2.95)



- 1) $\int n dl/l = n_{GW}$ P_{fus} = 2.7 GW
- 2) $\int n \, dl/l = n_{GW} P_{fus}$ = 2.7 GW
- 3) $\int n \, dl/l = 1.4 n_{GW}$ $n(\rho=0.9) = n_{GW}$ $P_{fus} = 3.7 \, GW$



Increase of density → **ELM characteristic changes** → **H to L-transition** → MARFE → disruption



H-mode density limit



Increase of density \rightarrow ELM characteristic changes \rightarrow H to L-transition \rightarrow MARFE \rightarrow disruption In the end MHD event, locked mode and disruption:

m/n = 2/1 mode suggests loss of electrical conductivity and unstable current density profile



Core density can exceed Greenwald limit

Pellet fuelling allows central $n > n_{GW}$

Pedestal density stays below $\rm n_{GW}$



Core density can exceed Greenwald limit

Pellet fuelling allows central $n > n_{GW}$

Pedestal density stays below n_{GW}, modest confinement degradation

ELMs suppressed with perturbation coils, pellets for HFS fuelling



Lang et al., EPS 2011

Divertor gas pressure in ITER and DEMO

B2-EIRENE (SOLPS 4.3) modeling, (A.S. Kukushkin, H.D. Pacher)



The pressure is determined by

- upstream plasma pressure;
- magnetic configuration;
- plasma-neutrals interaction in divertor

Similar sub-divertor pressure 1-15 Pa in ITER and DEMO

Is it possible to increase it significantly?

Can inter-cassette gaps be used for pumping? He de-enrichment?

IDD

Summary & work plan



- Further verify that in tokamaks density limit is a pedestal limit / not a core limit
- Improve density scaling (scaling parameters not independent, 0D approach too simple)
- Apply methods to increase core density above Greenwald limit
 - Pellet fuelling is pellet fuelling applicable in a tokamak DEMO?
 - Investigate density peaking at low collisionality how far can we go on present day tokamaks?
- Radiation limit in stellarators
 - Impurity control is not an optimization criterion of W7-X see C. Beidler et al.
 - HDH in W7-AS (or SDC in LHD) is a starting point, but not understood how it extrapolates to e.g. W7-X
- Divertor & fuelling
 - High density / detachment required for staying below critical heat flux (≤ 5 MW/m²?) see L. Boccaccini, A. Kallenbach
 - High neutral pressure desired for "simple" pumping solution see C Day et al.
 - Are geometrical solutions (e.g. pumping close to strike points) feasible?
 - Balance between neutral pressure and upstream plasma pressure?